

UNIVERSITY OF SWAZILAND

FACULTY OF SCIENCE

DEPARTMENT OF PHYSICS.

MAIN EXAMINATION 2006/2007

TITLE OF PAPER: STATISTICAL PHYSICS & THERMODYNAMICS

COURSE NUMBER: P 461

TIME ALLOWED : THREE HOURS

THIS PAPER CONTAINS FIVE QUESTIONS. ANSWER ANY **FOUR** QUESTIONS . ALL QUESTIONS CARRY EQUAL MARKS

THIS PAPER IS NOT TO BE OPENED UNTIL PERMISSION HAS BEEN GIVEN BY THE INVIGILATOR.

**QUESTION ONE.**

- (a) (i) Explain the difference between a **macrostate** and a **microstate** of a system of particles. (2 marks)
- (ii) What is meant by statistical weight of a macrostate in a system of particles? (2 marks)
- (iii) A classical system has 5 particles to be arranged in 3 energy levels. Using appropriate expression find the weights and hence the most probable distribution of these particles for the following cases:
1. The energy levels are non-degenerate.
  2. The energy levels are doubly degenerate (12 marks)
- (b) (i) What is meant by Phase space? (2 marks)
- (ii) Derive expressions for the volume element in phase space in terms of
1. Momentum  $p$
  2. Energy  $\epsilon$ . (3 + 4 marks)

**QUESTION TWO.**

- (a) Derive the Fermi-Dirac distribution function for a system of fermions,

$$n_S = \frac{g_S}{e^{-(\alpha + \beta \epsilon_S)} + 1}, \text{ where symbols have their usual meanings}$$

(12 marks)

- (b) (i) Given that the density of states for fermions is:

$$g(\epsilon)d\epsilon = \frac{4\pi V}{h^3} (2m)^{3/2} \epsilon^{1/2} d\epsilon$$

where symbols have their usual meanings, show that the Fermi energy of a system of fermions:

$$\epsilon_F = \frac{h^2}{2m} \left( \frac{3N}{8\pi V} \right)^{2/3}$$

(8 marks)

- (ii) Calculate the Fermi energy of a metal having density
- $8.5 \times 10^2 \text{ kg m}^{-3}$
- and atomic weight 40. (5 marks)

**QUESTION THREE.**

- (a) The Maxwell- Boltzmann distribution function for a system of classical particles is given by:

$$n_s = g_s e^{\alpha + \beta \epsilon_s},$$

where the symbols have their usual meanings. Such a system has 2000 particles distributed in three non-degenerate energy levels having energies 1 unit and 2 units and 3 units each. The total energy is 2600 units. Use the above distribution function to obtain the values of  $\alpha$  and  $\beta$  of this system and hence find its probable configuration. Verify your answer numerically.

(15 marks)

- (b) The differential form of Maxwell-Boltzmann distribution function in terms of the velocity  $v$  of the particles is given as:

$$n(v)dv = 4\pi N \left( \frac{m}{2\pi kT} \right)^{3/2} e^{-mv^2/2kT} v^2 dv$$

Use this expression to obtain:

- (i) The mean velocity
- (ii) The most probable velocity of the molecules of a classical gas.

[Note: see appendix for definite integrals ]

(10 marks)

**QUESTION FOUR.**

- (a) Use the Bose-Einstein distribution function of an assembly of identical non-interacting particles in thermal equilibrium to derive the Planck's radiation law for spectral distribution of energy radiated from a constant temperature enclosure. (9 marks)
- (b) Obtain an expression for the total energy per unit volume emitted from the enclosure at temperature T. (5 marks)
- (c) (i) State briefly what is *Bose-Einstein condensation*. (3 marks)
- (ii) The density of ideal gas consisting of particles having mass  $6.65 \times 10^{-27}$  kg is  $1.17 \times 10^{26} \text{ m}^{-3}$ .
1. Calculate the Bose temperature  $T_B$  of the gas. (5 marks)
  2. What fraction of the particles will be in the ground state at a temperature of  $0.1T_B$ . (3 marks)

$$\text{Given: } N = 2.612V \left( \frac{2\pi mkT_B}{h^2} \right)^{3/2}$$

**QUESTION FIVE.**

- (a) Derive the partition function of a classical gas:

$$Z = \frac{v}{h^3} (2\pi mkT)^{3/2}$$

(8 marks)

- (b) Show that the pressure of the classical gas:

$$P = NkT \frac{\partial \ln Z}{\partial V}$$

Hence derive the ideal gas equation  $P V = N k T$

(10 marks)

- (c) Calculate the translational partition function of an hydrogen molecule confined to a volume of  $100 \text{ cm}^3$  at  $300 \text{ K}$ .

(7 marks)

**Appendix 1**

Various definite integrals.

$$\int_0^{\infty} e^{-ax^2} dx = \frac{1}{2} \sqrt{\frac{\pi}{a}}$$

$$\int_0^{\infty} e^{-ax^2} x dx = \frac{1}{2a}$$

$$\int_0^{\infty} e^{-ax^2} x^3 dx = \frac{1}{2a^2}$$

$$\int_0^{\infty} e^{-ax^2} x^2 dx = \frac{1}{4} \sqrt{\frac{\pi}{a^3}}$$

$$\int_0^{\infty} e^{-ax^2} x^4 dx = \frac{3}{8a^2} \left(\frac{\pi}{a}\right)^{1/2}$$

$$\int_0^{\infty} e^{-ax^2} x^5 dx = \frac{1}{a^3}$$

$$\int_0^{\infty} \frac{x^3 dx}{e^x - 1} = \frac{\pi^4}{15}$$

$$\int_0^{\infty} x^{1/2} e^{-\lambda x} dx = \frac{\pi^{1/2}}{2\lambda^{3/2}}$$

$$\int_0^{\infty} \frac{x^4 e^x}{(e^x - 1)^2} dx = \frac{4\pi^4}{15}$$

$$\int_0^{\infty} \frac{x^{1/2}}{e^x - 1} dx = \frac{2.61\pi^{1/2}}{2}$$

**Appendix 2**

## Physical Constants.

<i>Quantity</i>	<i>symbol</i>	<i>value</i>
Speed of light	$c$	$3.00 \times 10^8 \text{ ms}^{-1}$
Plank's constant	$h$	$6.63 \times 10^{-34} \text{ J.s}$
Boltzmann constant	$k$	$1.38 \times 10^{-23} \text{ JK}^{-1}$
Electronic charge	$e$	$1.61 \times 10^{-19} \text{ C}$
Mass of electron	$m_e$	$9.11 \times 10^{-31} \text{ kg}$
Mass of proton	$m_p$	$1.67 \times 10^{-27} \text{ kg}$
Gas constant	$R$	$8.31 \text{ J mol}^{-1} \text{ K}^{-1}$
Avogadro's number	$N_A$	$6.02 \times 10^{23}$
Bohr magneton	$\mu_B$	$9.27 \times 10^{-24} \text{ JT}^{-1}$
Permeability of free space	$\mu_0$	$4\pi \times 10^{-7} \text{ Hm}^{-1}$
Stefan constant	$\sigma$	$5.67 \times 10^{-8} \text{ Wm}^{-2}\text{K}^{-4}$
Atmospheric pressure		$1.01 \times 10^5 \text{ Nm}^{-2}$
Mass of ${}_2^4\text{He}$ atom		$6.65 \times 10^{-27} \text{ kg}$
Mass of ${}_2^3\text{He}$ atom		$5.11 \times 10^{-27} \text{ kg}$
Volume of an ideal gas at STP		$22.4 \text{ l mol}^{-1}$