

UNIVERSITY OF SWAZILAND

FACULTY OF SCIENCE AND ENGINEERING

DEPARTMENT OF PHYSICS

SUPPLEMENTARY EXAMINATION 2015/2016

TITLE OF PAPER : CLASSICAL MECHANICS

COURSE NUMBER : P320

TIME ALLOWED : THREE HOURS

**INSTRUCTIONS : ANSWER ANY FOUR OUT OF FIVE
QUESTIONS.
EACH QUESTION CARRIES 25
MARKS.
MARKS FOR DIFFERENT SECTIONS
ARE SHOWN IN THE RIGHT-HAND
MARGIN.**

THIS PAPER HAS TWELVE PAGES, INCLUDING THIS PAGE.

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GIVEN BY THE INVIGILATOR.**

P320 CLASSICAL MECHANICS

Question one

- (a) For a certain dynamical system the kinetic energy T and potential energy V are given by

$$T = \frac{1}{2} (\dot{q}_1^2 + \dot{q}_1 \dot{q}_2 + \dot{q}_2^2)$$

$$V = \frac{3}{2} q_2^2$$

where q_1 , q_2 are the generalized coordinates.

- (i) Write down Lagrange's equations of motion and show that

$$\begin{cases} \ddot{q}_1 + \frac{1}{2} \ddot{q}_2 = 0 \\ \frac{1}{2} \ddot{q}_1 + \ddot{q}_2 = -3 q_2 \end{cases}$$

(4 marks)

- (ii) Eliminate \ddot{q}_1 from the Lagrange's equations of motion in (a)(i) to obtain a differential equation for q_2 and then show that the general solution of q_2 can be written as $q_2 = k_1 \cos(2t) + k_2 \sin(2t)$ where k_1 and k_2 are two arbitrary constants.

(4 marks)

- (iii) Substitute the general solution of q_2 in (a)(ii) back to any one equation in (a)(i) and integrate twice to show that the general solution of q_1 can be written as

$$q_1 = -\frac{k_1}{2} \cos(2t) - \frac{k_2}{2} \sin(2t) + k_3 t + k_4 \quad \text{where } k_3 \text{ and } k_4 \text{ are another two arbitrary constants.}$$

(4 marks)

- (iv) Write down their canonical momenta p_{q_1} & p_{q_2} and then deduce that

$$\begin{cases} \dot{q}_1 = \frac{4}{3} p_{q_1} - \frac{2}{3} p_{q_2} \\ \dot{q}_2 = -\frac{2}{3} p_{q_1} + \frac{4}{3} p_{q_2} \end{cases}$$

(5 marks)

- (b) (i) Use $H = T + V$ and the results in (a)(iv) to write down the Hamiltonian of the system and deduce that

$$H = \frac{2}{3} (p_{q_1}^2 + p_{q_2}^2 - p_{q_1} p_{q_2}) + \frac{3}{2} q_2^2$$

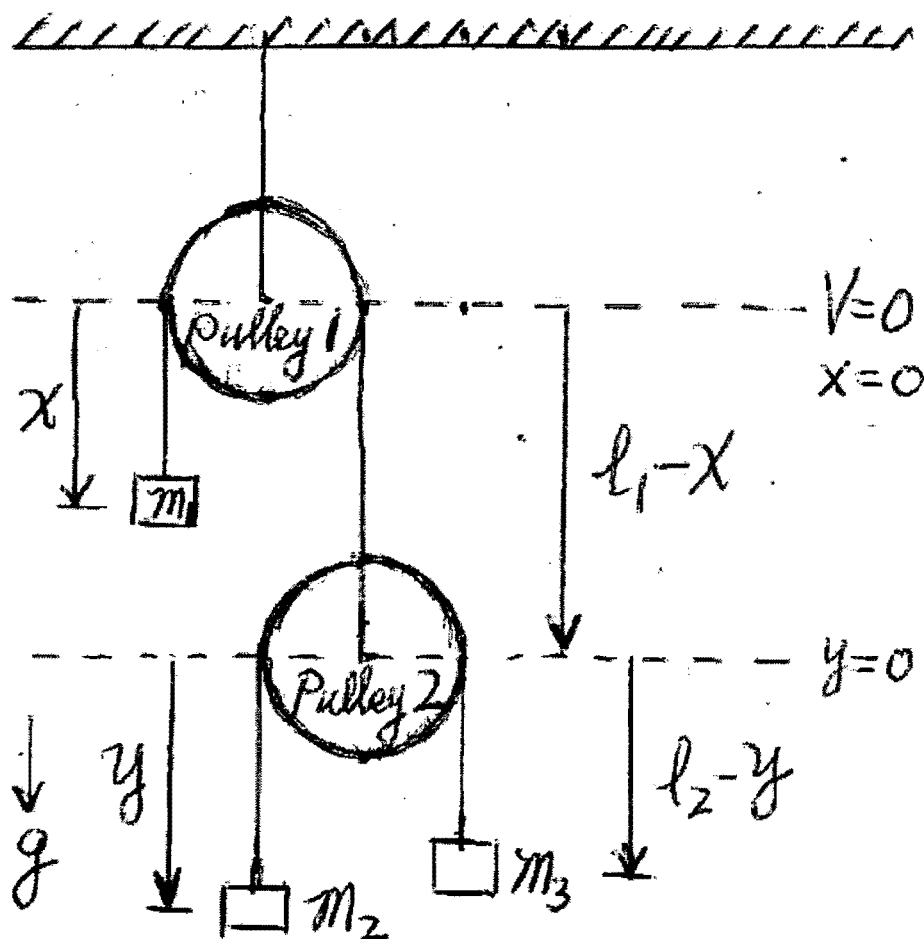
(4 marks)

- (ii) From the Hamiltonian in (b)(i), write down the equations of motion of the system.

(4 marks)

Question two

- (a) Consider the double pulley system as shown in the following diagram :



Assume the pulleys are massless and let l_1 & l_2 be the length of rope hanging freely from *pulley 1* and *pulley 2* respectively. Assume the pulley system subject only to gravitational force with zero gravitational potential set at $x = 0$.

- (i) Write down the Lagrangian L for the given system and show that it can be simplified to the following expression :

$$L = \frac{1}{2} (m_1 + m_2 + m_3) \dot{x}^2 + \frac{1}{2} (m_2 + m_3) \dot{y}^2 + (m_3 - m_2) \dot{x} \dot{y} + m_1 g x + m_2 g (l_1 - x + y) + m_3 g (l_1 - x + l_2 - y)$$

(8 marks)

- (ii) Write down Lagrange's equations of motion and show that

$$\begin{cases} (m_1 + m_2 + m_3) \ddot{x} + (m_3 - m_2) \ddot{y} = (m_1 - m_2 - m_3) g \\ (m_3 - m_2) \ddot{x} + (m_3 + m_2) \ddot{y} = (m_2 - m_3) g \end{cases}$$

(4 marks)

Question two (continued)

(b) If H denotes the Hamiltonian function and L is the Lagrangian function, use the definition $H = \sum_{\alpha=1}^n p_{\alpha} \dot{q}_{\alpha} - L$ (where p_{α} and q_{α} ($\alpha=1,2,\dots,n$) are the generalized momenta and coordinates respectively, i.e., $H = H(q_1, \dots, q_n, p_1, \dots, p_n, t)$,

$L = L(q_1, \dots, q_n, \dot{q}_1, \dots, \dot{q}_n, t)$, $p_{\alpha} = \frac{\partial L}{\partial \dot{q}_{\alpha}}$ and $\dot{p}_{\alpha} = \frac{\partial L}{\partial q_{\alpha}}$) to show that

(i) $\dot{q}_{\alpha} = \frac{\partial H}{\partial p_{\alpha}} \quad \alpha = 1, 2, \dots, n$ (4 marks)

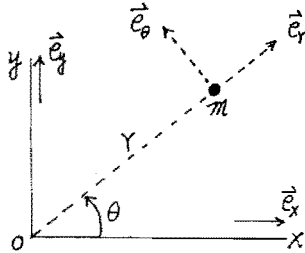
(ii) $\dot{p}_{\alpha} = -\frac{\partial H}{\partial q_{\alpha}} \quad \alpha = 1, 2, \dots, n$ (4 marks)

(iii) $\frac{dF}{dt} = [F, H] + \frac{\partial F}{\partial t}$ where $F = F(q_1, \dots, q_n, p_1, \dots, p_n, t)$ and

$$[F, H] \equiv \sum_{\alpha=1}^n \left(\frac{\partial F}{\partial q_{\alpha}} \frac{\partial H}{\partial p_{\alpha}} - \frac{\partial F}{\partial p_{\alpha}} \frac{\partial H}{\partial q_{\alpha}} \right) \quad (5 marks)$$

Question three

Consider a particle of mass m acted on by an attractive central force of $\vec{F} = -\vec{e}_r \frac{k}{r^n}$, where k is a positive constant and $n > 1$, and moving in a 2-D plane described by the plane polar coordinates as shown in the diagram below.



The kinetic energy of this particle in this plane polar coordinate is

$$T = \frac{m}{2} (\dot{r}^2 + r^2 \dot{\theta}^2) \dots\dots (1)$$

- (i) From $V = - \int_{r_0}^r \vec{F} \cdot d\vec{l}$ where $d\vec{l} = d\vec{r} = \vec{e}_r dr + \vec{e}_\theta r d\theta$ & $r_0 \rightarrow \infty$, find the potential energy V of this particle in this plane polar coordinate under the given force $\vec{F} = -\vec{e}_r \frac{k}{r^n}$ where k is a constant. Show that

$$V = - \left(\frac{k}{(n-1)r^{n-1}} \right) \dots\dots (2) \quad (3 \text{ marks})$$

- (ii) Write down the Lagrange equations of motion for this system and show that

$$\begin{cases} m\ddot{r} = \left(m r \dot{\theta}^2 - \frac{k}{r^n} \right) \dots\dots (3) \\ \frac{d}{dt} (m r^2 \dot{\theta}) = 0 \dots\dots (4) \end{cases} \quad (5 \text{ marks})$$

- (iii) Write its (r, θ) respective momentums, i.e., p_r & p_θ , and show that the angular momentum p_θ is a constant and then eq.(1) & eq.(3) can be rewritten in terms of this constant p_θ as $T = \frac{m \dot{r}^2}{2} + \frac{p_\theta^2}{2 m r^2} \dots\dots (1)'$ and $m\ddot{r} = \left(\frac{p_\theta^2}{m r^3} - \frac{k}{r^n} \right) \dots\dots (3)'$

respectively. (4 mark)

- (iv) Multiply eq.(1)' by dr and show that it can be rewritten as $d(T + V) = 0$ and thus this implies the total energy $(T + V)$ is also a constant. (5 marks)

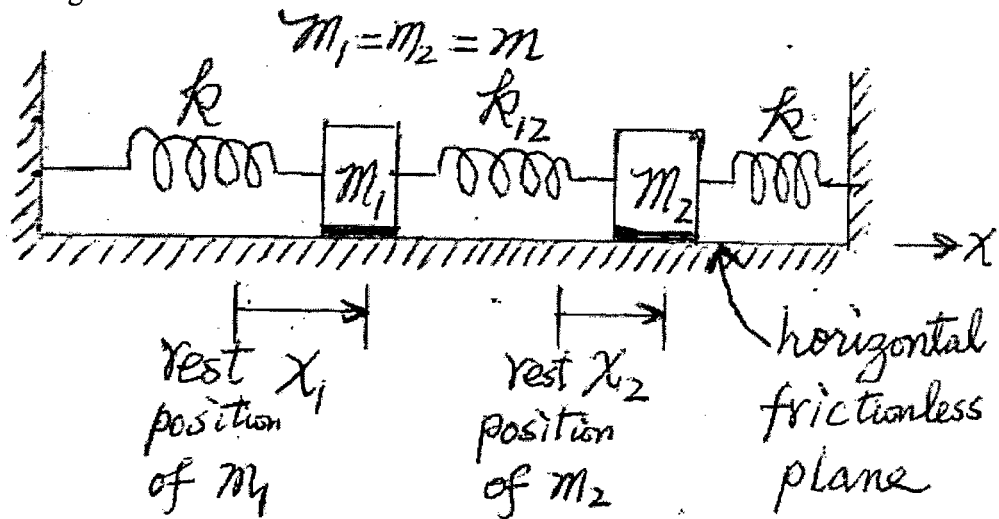
(Hint : $\ddot{r} dr = \frac{d\dot{r}}{dt} dr = d\dot{r} \frac{dr}{dt} = \dot{r} d\dot{r} = d\left(\frac{\dot{r}^2}{2}\right)$)

- (v) For circular orbits, i.e., r is a constant, find a relation between the kinetic and potential energies and show that $T = -\frac{n-1}{2} V$. (8 marks)

(Hint : Use eq.(1)', eq.(3)', eq.(2) and $\dot{r} = 0$ & $\ddot{r} = 0$)

Question four

- (a) Two identical simple harmonic oscillators of mass m and spring constant k are joined by a spring of spring constant k_{12} and allowed to oscillate on a horizontal frictionless plane along x - direction as shown below :



where x_1 & x_2 are the displaced lengths from the rest positions of m_1 & m_2 respectively. The Lagrangian for the system can be written as:

$$L = \frac{1}{2} m \dot{x}_1^2 + \frac{1}{2} m \dot{x}_2^2 - \frac{1}{2} k x_1^2 - \frac{1}{2} k_{12} (x_1 - x_2)^2 - \frac{1}{2} k x_2^2$$

- (i) Write down the equations of motion and deduce that

$$\begin{cases} \ddot{x}_1 = -\left(\frac{k + k_{12}}{m}\right) x_1 + \left(\frac{k_{12}}{m}\right) x_2 \\ \ddot{x}_2 = \left(\frac{k_{12}}{m}\right) x_1 - \left(\frac{k + k_{12}}{m}\right) x_2 \end{cases}$$

(4 marks)

- (ii) Set $x_1 = \hat{X}_1 e^{i\omega t}$ and $x_2 = \hat{X}_2 e^{i\omega t}$ (where \hat{X}_1 and \hat{X}_2 are constants) and deduce from the equations in (a)(i) the matrix equation $-\omega^2 X = A X$ where

$$X = \begin{pmatrix} \hat{X}_1 \\ \hat{X}_2 \end{pmatrix} \text{ and } A = \begin{pmatrix} -\left(\frac{k + k_{12}}{m}\right) & \left(\frac{k_{12}}{m}\right) \\ \left(\frac{k_{12}}{m}\right) & -\left(\frac{k + k_{12}}{m}\right) \end{pmatrix}$$

(4 marks)

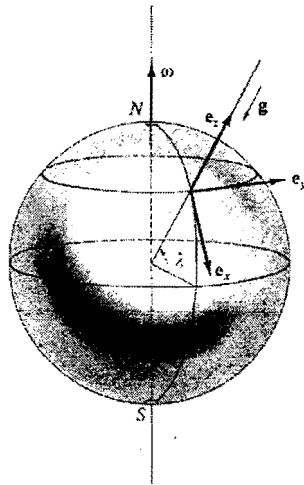
- (iii) Show that the eigenfrequencies ω of this coupled system are

$$\omega_1 = \sqrt{\frac{k}{m}} \quad \& \quad \omega_2 = \sqrt{\frac{k + 2k_{12}}{m}}$$

(4 marks)

Question four (continued)

(b)



If a particle is projected vertically upward with an initial speed v_0 to a height h above a point on the earth's surface at northern latitude λ , show that it strikes the ground at a point $\frac{4}{3} \omega \cos(\lambda) \sqrt{\frac{8 h^3}{g}}$ to the west of the initial throwing point. Neglect air resistance and only consider small vertical height. (13 marks)

(Hint :

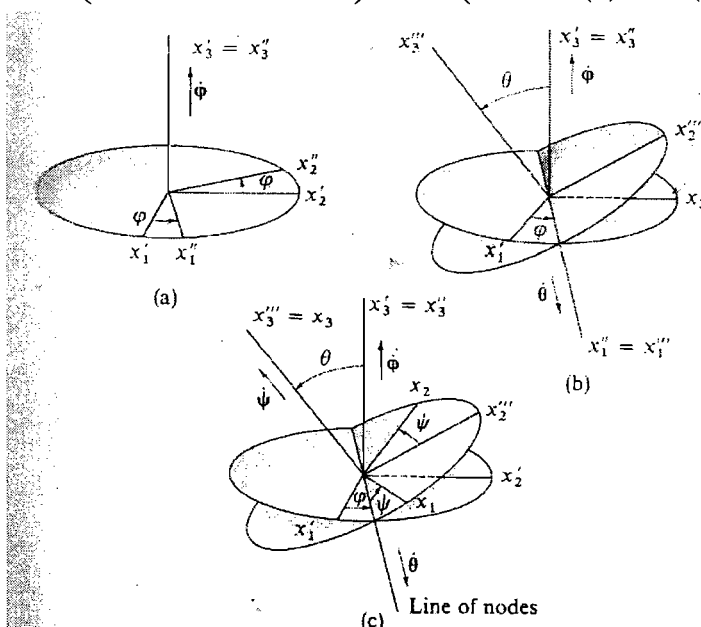
$$\vec{a}_{eff} \approx \vec{e}_z (-g) - 2 \vec{\omega} \times \vec{v}_r, \quad \vec{v}_r \approx \vec{e}_z (v_0 - g t), \quad \vec{\omega} = \vec{e}_x (-\omega \cos(\lambda)) + \vec{e}_z (\omega \sin(\lambda))$$

$$\text{and } v_0 = \sqrt{2 g h}, \quad (\text{total time for the given motion}) = \frac{2 v_0}{g})$$

Question five

- (a) The fixed (or inertia) coordinate system X' shares the same origin with the body coordinate system X such that only rotational motion is considered. The rotational velocity $\vec{\omega}$ of the body system with respect to the fixed system are breaking down into three independent angular velocities, i.e., $\vec{\omega} = \vec{\dot{\phi}} + \vec{\dot{\theta}} + \vec{\dot{\psi}}$ where (ϕ, θ, ψ) are Euler's angles. We use two intermediate coordinate systems X'' & X''' to bridge between X' & X systems such that $X'' = \lambda_\phi X'$, $X''' = \lambda_\theta X''$ & $X = \lambda_\psi X'''$ as shown in the following diagram, where

$$\lambda_\phi = \begin{pmatrix} \cos(\phi) & \sin(\phi) & 0 \\ -\sin(\phi) & \cos(\phi) & 0 \\ 0 & 0 & 1 \end{pmatrix}, \lambda_\theta = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos(\theta) & \sin(\theta) \\ 0 & -\sin(\theta) & \cos(\theta) \end{pmatrix}, \lambda_\psi = \begin{pmatrix} \cos(\psi) & \sin(\psi) & 0 \\ -\sin(\psi) & \cos(\psi) & 0 \\ 0 & 0 & 1 \end{pmatrix}$$



- (i) Since the direction of $\vec{\dot{\phi}}$ is along x_3' -axis (which is the same as x_3'' -axis) with

the magnitude of $\dot{\phi}$ thus $(\vec{\dot{\phi}})'' = \begin{pmatrix} 0 \\ 0 \\ \dot{\phi} \end{pmatrix}$ in X'' system, show that $\vec{\dot{\phi}}$ in X

system (i.e., the body system) is $(\vec{\dot{\phi}}) = \begin{pmatrix} \dot{\phi} \sin(\theta) \sin(\psi) \\ \dot{\phi} \sin(\theta) \cos(\psi) \\ \dot{\phi} \cos(\theta) \end{pmatrix}$ in X system. In other

words, show that

$$\begin{pmatrix} \dot{\phi} \sin(\theta) \sin(\psi) \\ \dot{\phi} \sin(\theta) \cos(\psi) \\ \dot{\phi} \cos(\theta) \end{pmatrix} = \lambda_\psi \lambda_\theta \begin{pmatrix} 0 \\ 0 \\ \dot{\phi} \end{pmatrix}$$

(6 marks)

Question five (continued)

(ii) Since the direction of $\vec{\theta}$ is along x_1'' -axis (which is the same as x_1''' -axis)

with the magnitude of $\dot{\theta}$ thus $(\vec{\theta})''' = \begin{pmatrix} \dot{\theta} \\ 0 \\ 0 \end{pmatrix}$ in X''' system, show that $\vec{\theta}$ in X

system (i.e., the body system) is $(\vec{\theta}) = \begin{pmatrix} \dot{\theta} \cos(\psi) \\ -\dot{\theta} \sin(\psi) \\ 0 \end{pmatrix}$ in X system. In other

words, show that $\begin{pmatrix} \dot{\theta} \cos(\psi) \\ -\dot{\theta} \sin(\psi) \\ 0 \end{pmatrix} = \lambda_\psi \begin{pmatrix} \dot{\theta} \\ 0 \\ 0 \end{pmatrix}$. **(4 marks)**

(Note: Since the direction of $\vec{\psi}$ is along x_3''' -axis (which is the same as x_3

-axis) with the magnitude of $\dot{\psi}$ thus $(\vec{\psi}) = \begin{pmatrix} 0 \\ 0 \\ \dot{\psi} \end{pmatrix}$ in X system. Then the

rotational velocity $\vec{\omega} = \vec{\phi} + \vec{\theta} + \vec{\psi}$ in X system (i.e., body system) can be

written in terms of Euler's angles as $(\vec{\omega}) = \begin{pmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{pmatrix} = \begin{pmatrix} \dot{\phi} \sin(\theta) \sin(\psi) + \dot{\theta} \cos(\psi) \\ \dot{\phi} \sin(\theta) \cos(\psi) - \dot{\theta} \sin(\psi) \\ \dot{\phi} \cos(\theta) + \dot{\psi} \end{pmatrix}$)

(b) (i) By proper choice of the orientation of the body coordinate system, the inertia tensor I (i.e., rotational mass) of a rigid body can be in the form of a diagonal matrix,

i.e., $I = \begin{pmatrix} I_1 & 0 & 0 \\ 0 & I_2 & 0 \\ 0 & 0 & I_3 \end{pmatrix}$, thus its rotational kinetic energy is

$$T_{rot} = \frac{1}{2} I_1 \omega_1^2 + \frac{1}{2} I_2 \omega_2^2 + \frac{1}{2} I_3 \omega_3^2.$$

Consider a pure rotational motion of the rigid body under no external force, then its

Lagrangian is $L = T_{rot} = \frac{1}{2} I_1 \omega_1^2 + \frac{1}{2} I_2 \omega_2^2 + \frac{1}{2} I_3 \omega_3^2$ where

ω_1 , ω_2 & ω_3 are functions of $(\phi, \theta, \psi, \dot{\phi}, \dot{\theta}, \dot{\psi})$ as those given in (a)(ii).

Write down the Lagrange equation of motion for ψ , i.e., $\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{\psi}} \right) = \frac{\partial L}{\partial \psi}$, and

show that it can be simplified to

$$(I_1 - I_2) \omega_1 \omega_2 - I_3 \dot{\omega}_3 = 0.$$

(13 marks)

Question five (continued again)

- (ii) From $(I_1 - I_2)\omega_1 \omega_2 - I_3 \dot{\omega}_3 = 0$ derived in (b)(i), the other two independent equations for the above rigid body's motion can be simply written down directly as $(I_2 - I_3)\omega_2 \omega_3 - I_1 \dot{\omega}_1 = 0$ & $(I_3 - I_1)\omega_3 \omega_1 - I_2 \dot{\omega}_2 = 0$ without going through similar derivations as done in (b)(i). Explain briefly why so ?
(2 marks)

Useful informations

$$V = - \int \vec{F} \cdot d\vec{l} \quad \text{and reversely} \quad \vec{F} = -\vec{\nabla} V$$

$$L = T - V = L(q_1, q_2, \dots, q_n, \dot{q}_1, \dot{q}_2, \dots, \dot{q}_n, t)$$

$$p_\alpha = \frac{\partial L}{\partial \dot{q}_\alpha} \quad \text{and} \quad \dot{p}_\alpha = \frac{\partial L}{\partial q_\alpha}$$

$$H = \sum_{\alpha=1}^n (p_\alpha \dot{q}_\alpha) - L = H(q_1, q_2, \dots, q_n, \dot{q}_1, \dot{q}_2, \dots, \dot{q}_n, t)$$

$$\dot{q}_\alpha = \frac{\partial H}{\partial p_\alpha} \quad \text{and} \quad \dot{p}_\alpha = -\frac{\partial H}{\partial q_\alpha}$$

$$[u, v] \equiv \sum_{\alpha=1}^n \left(\frac{\partial u}{\partial q_\alpha} \frac{\partial v}{\partial p_\alpha} - \frac{\partial u}{\partial p_\alpha} \frac{\partial v}{\partial q_\alpha} \right)$$

$$G = 6.673 \times 10^{-11} \frac{N m^2}{kg^2}$$

$$\text{radius of earth } r_E = 6.4 \times 10^6 \text{ m}$$

$$\text{mass of earth } m_E = 6 \times 10^{24} \text{ kg}$$

$$\text{earth attractive potential} \equiv -\frac{k}{r} \quad \text{where} \quad k = G m m_E$$

$$\varepsilon = \sqrt{1 + \frac{2 E l^2}{\mu k^2}} \quad \{(\varepsilon = 0, \text{ circle}), (0 < \varepsilon < 1, \text{ ellipse}), (\varepsilon = 1, \text{ parabola}), \dots\}$$

$$\mu = \frac{m_1 m_2}{m_1 + m_2} \approx m_1 \quad \text{if} \quad m_2 \gg m_1$$

$$\text{For elliptical orbit, i.e., } 0 < \varepsilon < 1, \text{ then} \left\{ \begin{array}{l} \text{semi-major } a = \frac{k}{2|E|} \\ \text{semi-minor } b = \frac{l}{\sqrt{2\mu|E|}} \\ \text{period } \tau = \frac{2\mu}{l} (\pi a b) \\ r_{\min} = a(1 - \varepsilon) \quad \& \quad r_{\max} = a(1 + \varepsilon) \end{array} \right.$$

for plane polar (r, θ) system with unit vectors $(\vec{e}_r, \vec{e}_\theta)$, we have

$$\left\{ \begin{array}{l} \vec{v} = \vec{e}_r \dot{r} + \vec{e}_\theta r \dot{\theta} \\ \vec{a} = \vec{e}_r (\ddot{r} - r \dot{\theta}^2) + \vec{e}_\theta (2\dot{r} \dot{\theta} + r \ddot{\theta}) \end{array} \right.$$

$$\vec{\nabla} f = \vec{e}_r \frac{\partial f}{\partial r} + \vec{e}_\theta \frac{1}{r} \frac{\partial f}{\partial \theta}$$

Useful informations (continued)

$$I = \begin{pmatrix} \sum_{\alpha} m_{\alpha} (x_{\alpha,2}^2 + x_{\alpha,3}^2) & -\sum_{\alpha} m_{\alpha} x_{\alpha,1} x_{\alpha,2} & -\sum_{\alpha} m_{\alpha} x_{\alpha,1} x_{\alpha,3} \\ -\sum_{\alpha} m_{\alpha} x_{\alpha,2} x_{\alpha,1} & \sum_{\alpha} m_{\alpha} (x_{\alpha,1}^2 + x_{\alpha,3}^2) & -\sum_{\alpha} m_{\alpha} x_{\alpha,2} x_{\alpha,3} \\ -\sum_{\alpha} m_{\alpha} x_{\alpha,3} x_{\alpha,1} & -\sum_{\alpha} m_{\alpha} x_{\alpha,3} x_{\alpha,2} & \sum_{\alpha} m_{\alpha} (x_{\alpha,1}^2 + x_{\alpha,2}^2) \end{pmatrix}$$

$$\vec{F}_{\text{eff}} = \vec{F} - m \ddot{\vec{R}}_f - m \dot{\vec{\omega}} \times \vec{r} - m \vec{\omega} \times (\vec{\omega} \times \vec{r}) - 2 m \vec{\omega} \times \vec{v}_r \quad \text{where}$$

$$\vec{r}' = \vec{R} + \vec{r} \quad \text{and}$$

\vec{r}' refers to fixed (inertial system)

\vec{r} refers to rotatinal (non-inertial system) rotates with $\vec{\omega}$ to \vec{r}' system

\vec{R} from the origin of \vec{r}' to the origin of \vec{r}

$$\vec{v}_r = \left(\frac{d\vec{r}}{dt} \right)_r$$